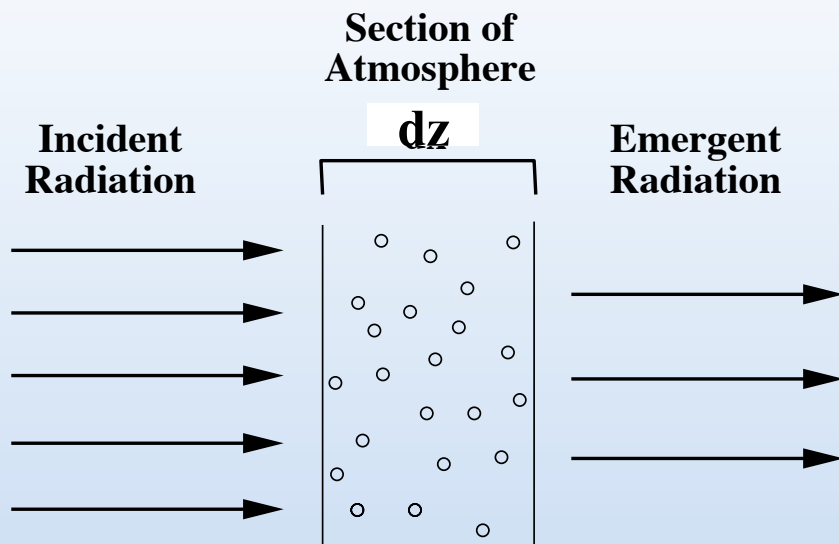


Atmospheric Chemistry

Fall 2014

Lecture 3

Absorption of Solar Radiation



Beer's Law Absorption:
reduction of incident intensity is proportional to the incident intensity, the density of scatterers and the thickness of the slab of atmosphere

$$dI = -\sigma \cdot n \cdot I \cdot dz$$

The proportionality constant is the "cross section", σ , with the units of cm^2

*Rearrange and integrate over z
(from z to infinity)*

$$\int \frac{dI}{I} = -\sigma \int n \cdot dz$$

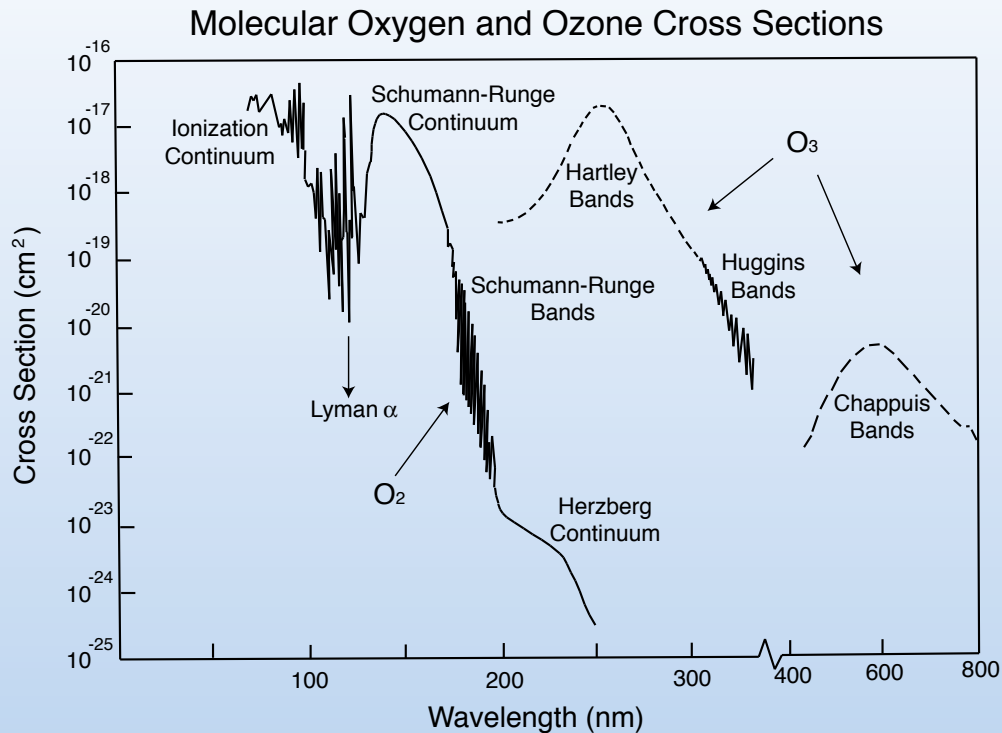
*Solve for Intensity as a function
of z (at a single wavelength)*

$$I(z, \lambda) = I_o(\lambda) \cdot e^{-N(z) \cdot \sigma(\lambda)}$$

*where N is the column density of
absorbers above the altitude z*

$$N(z) = \int_z^{\infty} n(z) dz$$

Oxygen and Ozone Cross Sections



*What does a cross section of 10^{-17} cm^2 mean?
Intensity of radiation will be reduced to $1/e$
when column density above altitude = 10^{17} cm^{-2} .*

$$I(z, \lambda) = I_o(\lambda) \cdot e^{-N(z) \cdot \sigma(\lambda)}$$

$$O_2(z) = O_{2o} e^{-(z-z_o)/H}$$

$$N_{O_2}(z_0) = \int_{z_0}^{\infty} n_{O_2} \cdot e^{-(z-z_0)/H} dz$$

$$N_{O_2}(z_0) = n_{O_2} \cdot H$$

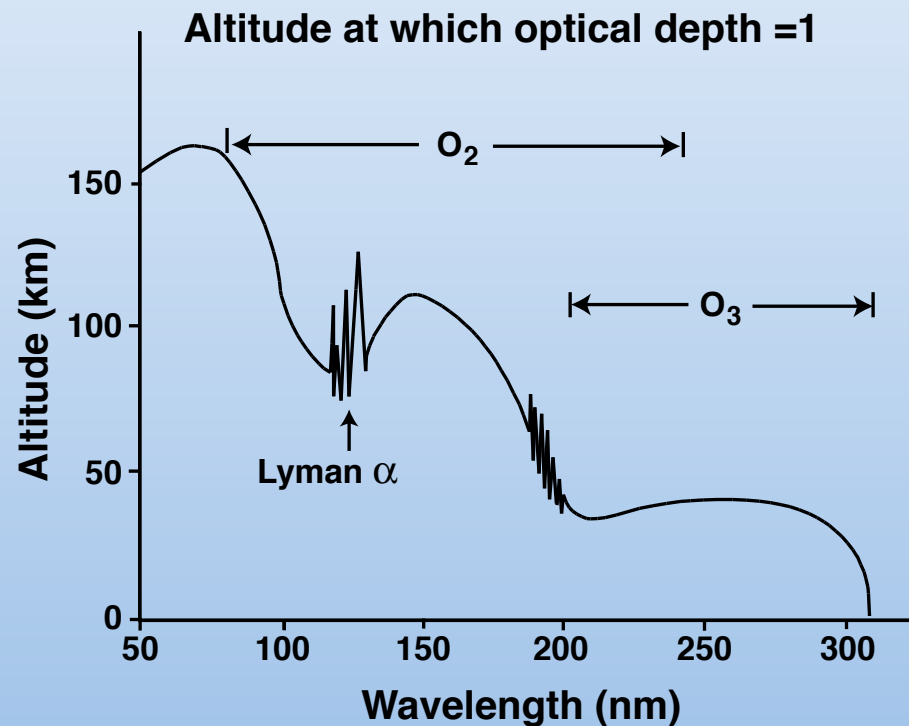
*Column density of O_2 is 10^{17}
when density is $10^{17}/H$ or
about $1.4 \times 10^{11} \text{ cm}^{-3}$.*

*O_2 density at surface is about
 5×10^{18} . Thus $1/e$ absorption is
at pressure of about .003 hPa –
well above stratosphere*

Penetration of Solar Radiation

$$I(z, \lambda) = I_o(\lambda) \cdot e^{-N(z) \cdot \sigma(\lambda)}$$

$$\tau = N \cdot \sigma \quad \textit{Optical Depth}$$



Photolysis of O₂ and O₃



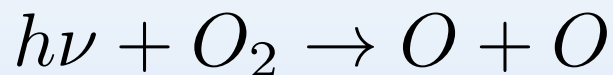
$$\frac{dn}{dt} = -J \cdot n$$

J has units of sec⁻¹

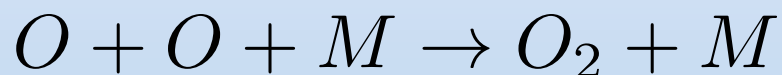
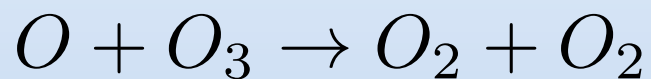
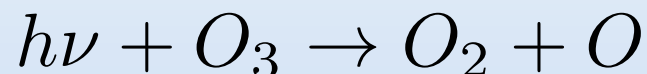
J is the rate of photon absorptions per second that lead to photolysis.

$$J(z) = \int \sigma(\lambda) \cdot I(\lambda, z) d\lambda$$

Chapman Mechanism for Formation of the Ozone Layer



Initiation



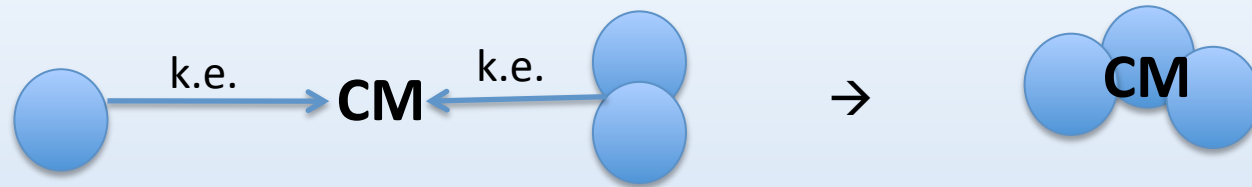
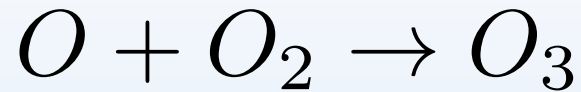
Propagation

Termination

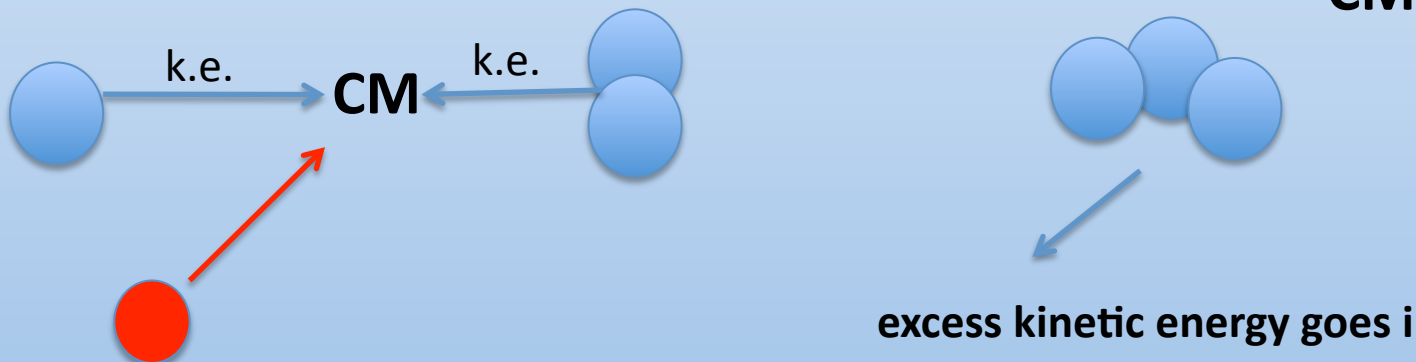
*Put forward by Sydney
Chapman in 1930*



Three-body reactions



kinetic energy must be absorbed by O_3 molecule into an excited state because CM is at rest



excess kinetic energy goes into relative velocities of particles away from CM

Time Constants

$$\tau = \frac{1}{J} \quad \textit{photolysis}$$

$$\tau = \frac{1}{k \cdot [n]} \quad \textit{two - body}$$

$$\tau = \frac{1}{k \cdot [n] \cdot [M]} \quad \textit{three - body}$$

Photodissociation Coefficients versus Altitude

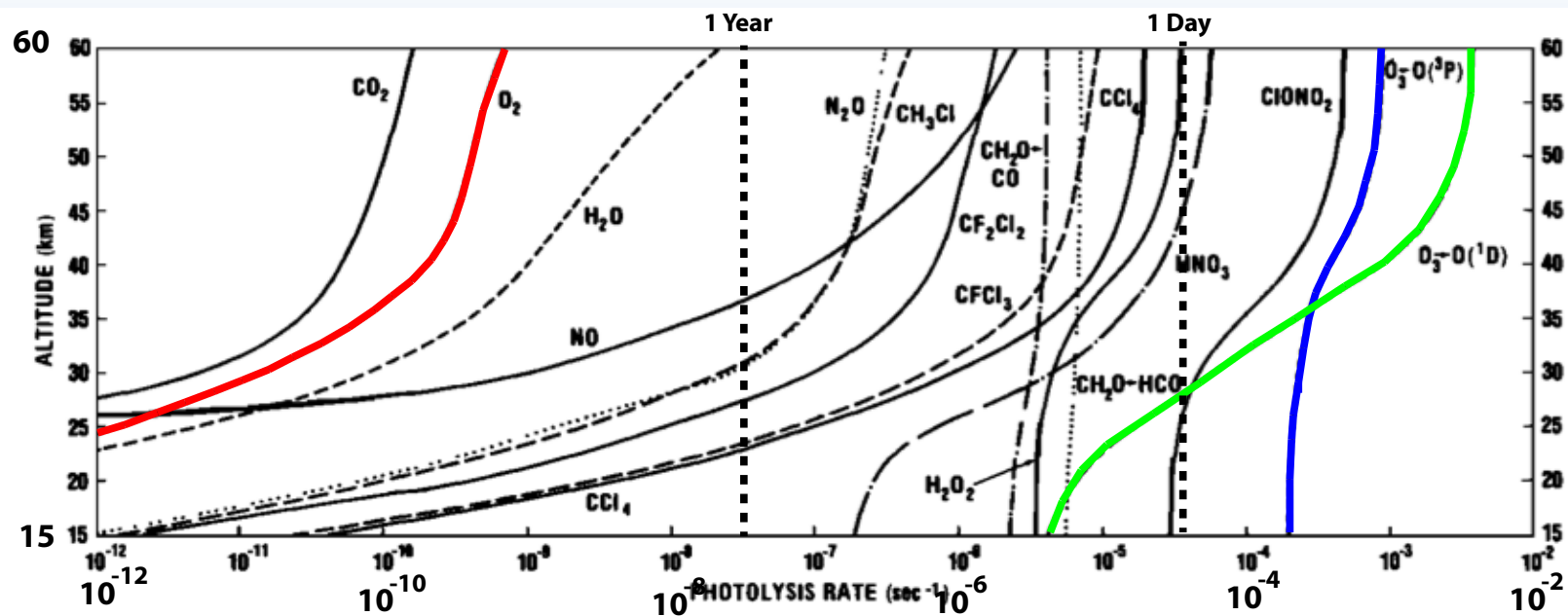
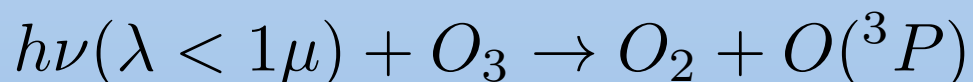
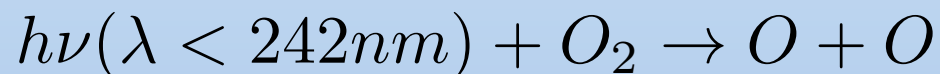
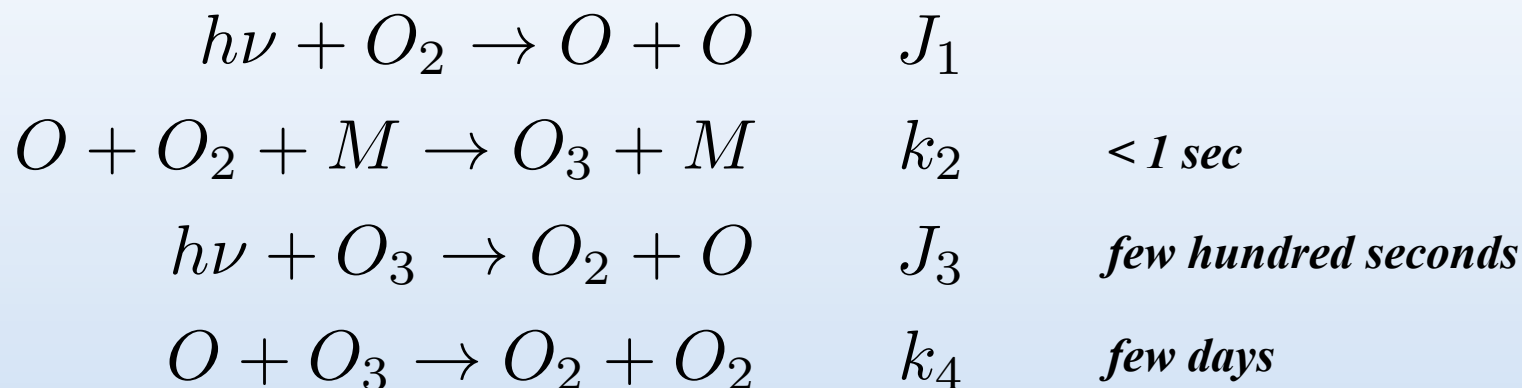


Fig. 1. Calculated 24-hour average J coefficients as a function of altitude. For $h\nu + \text{NO}_2 \rightarrow \text{NO} + \text{O}$, $h\nu + \text{NO}_3 \rightarrow \text{NO}_2 + \text{O}$, and $h\nu + \text{NO}_3 \rightarrow \text{NO} + \text{O}_2$, constant J coefficients were used. The values were 0.005 , 0.05 , and 0.02 s^{-1} , respectively. The dissociation of NO is discussed in the text.



Chapman Mechanism Equations

Time Scale at ~40 km



$$[O_x] = [O] + [O_3] \quad \text{OddOxygen}$$

$$\frac{d[O_x]}{dt} = 2 \cdot J_1 \cdot [O_2] - 2 \cdot k_4 \cdot \left(\frac{[O]}{[O_3]} \right) \cdot [O_3]^2$$

$$\frac{[O]}{[O_3]} = \frac{J_3}{k_2 \cdot [O_2] \cdot [M]}$$

$$[O_x] = \left(\frac{J_1 \cdot k_2 \cdot [M]}{k_4 \cdot J_3} \right)^{1/2} \cdot [O_2]$$

Reference Atmosphere

Altitude(km)	Temperature (K)	Air Density (cm^{-3})	Ozone (cm^{-3})
0	288	2.55×10^{19}	7.6×10^{11}
2	275	2.09×10^{19}	6.8×10^{11}
4	262	1.70×10^{19}	5.8×10^{11}
6	249	1.37×10^{19}	5.7×10^{11}
8	236	1.09×10^{19}	6.5×10^{11}
10	223	8.60×10^{18}	1.13×10^{12}
12	217	6.49×10^{18}	2.02×10^{12}
14	217	4.74×10^{18}	2.35×10^{12}
16	217	3.46×10^{18}	2.95×10^{12}
18	217	2.53×10^{18}	4.04×10^{12}
20	217	1.85×10^{18}	4.77×10^{12}
22	219	1.34×10^{18}	4.86×10^{12}
24	221	9.76×10^{17}	4.54×10^{12}
26	223	7.12×10^{17}	4.03×10^{12}
28	225	5.21×10^{17}	3.24×10^{12}
30	227	3.83×10^{17}	2.52×10^{12}
32	228	2.82×10^{17}	2.03×10^{12}
34	233	2.06×10^{17}	1.58×10^{12}
36	239	1.51×10^{17}	1.22×10^{12}
38	245	1.12×10^{17}	8.73×10^{11}
40	250	8.31×10^{16}	6.07×10^{11}
42	256	6.23×10^{16}	3.98×10^{11}
44	261	4.70×10^{16}	2.74×10^{11}
46	267	3.56×10^{16}	1.69×10^{11}
48	271	2.74×10^{16}	1.03×10^{11}
50	271	2.14×10^{16}	6.64×10^{10}
52	269	1.68×10^{16}	3.84×10^{10}
54	264	1.33×10^{16}	2.55×10^{10}
56	258	1.05×10^{16}	1.61×10^{10}
58	253	8.24×10^{15}	1.12×10^{10}
60	247	6.44×10^{15}	7.33×10^9

Some Important 2-Body Reaction Rate Coefficients

$$k(T) = A \cdot e^{-\frac{E}{R}/T}$$

Reaction	A-Factor	E/R	k(298K)
$O + O_3 \rightarrow O_2 + O_2$	8.0×10^{-12}	2060	8.0×10^{-15}
$O + NO_2 \rightarrow NO + O_2$	6.5×10^{-12}	-120	9.7×10^{-12}
$O_3 + NO \rightarrow NO_2 + O_2$	2.0×10^{-12}	1400	1.8×10^{-14}
$O + HO_2 \rightarrow OH + O_2$	3.0×10^{-11}	-200	5.9×10^{-11}
$O + OH \rightarrow H + O_2$	2.2×10^{-11}	-120	3.3×10^{-11}
$O_3 + HO_2 \rightarrow OH + O_2 + O_2$	1.1×10^{-14}	500	2.0×10^{-15}
$O_3 + OH \rightarrow HO_2 + O_2$	1.6×10^{-12}	940	6.8×10^{-14}
$Cl + O_3 \rightarrow ClO + O_2$	2.9×10^{-11}	260	1.2×10^{-11}
$ClO + O \rightarrow Cl + O_2$	3.0×10^{-11}	-70	3.8×10^{-11}
$Br + O_3 \rightarrow BrO + O_2$	1.7×10^{-11}	800	1.2×10^{-12}
$BrO + ClO \rightarrow Br + Cl + O_2$	2.9×10^{-12}	-220	6.1×10^{-12}
$BrO + BrO \rightarrow Br + Br + O_2$	1.4×10^{-12}	-150	2.3×10^{-12}
$Cl + CH_4 \rightarrow HCl + CH_3$	1.1×10^{-11}	1400	1.0×10^{-13}
$OH + HCl \rightarrow H_2O + Cl$	2.6×10^{-12}	350	8.0×10^{-13}

Some Important 3-Body Reaction Rate Coefficients

$$\text{LowPressureLimit} \quad k_o(T) = k_o^{300} \cdot \left(\frac{T}{300}\right)^{-n}$$

$$\text{HighPressureLimit} \quad k_\infty(T) = k_\infty^{300} \cdot \left(\frac{T}{300}\right)^{-m}$$

Reaction	k_o^{300}	n	k_∞^{300}	m
$O + O_2 + M \rightarrow O_3 + M$	6.0×10^{-34}	2.3	-	-
$OH + NO_2 + M \rightarrow HNO_3 + M$	2.6×10^{-30}	3.2	2.4×10^{-11}	1.3
$H + O_2 + M \rightarrow HO_2 + M$	5.7×10^{-32}	1.6	7.5×10^{-11}	0
$ClO + NO_2 + M \rightarrow ClONO_2 + M$	1.8×10^{-31}	3.4	1.5×10^{-11}	1.9