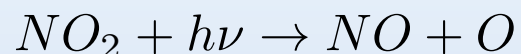
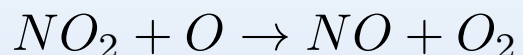
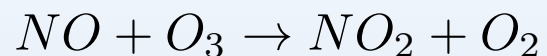


Atmospheric Chemistry

Lecture 5

Rate Determining Step for NO_x Catalysis of Ozone Loss



*Rapid reactions set
relationship between NO
and NO₂ in the stratosphere*

NO continuity equation

$$\frac{dNO}{dt} = -k_{NO,O_3} \cdot [O_3] \cdot [NO] + k_{O,NO_2} \cdot [O] \cdot [NO_2] + J_{NO_2} \cdot [NO_2]$$

$$\frac{NO}{NO_2} = \frac{J_{NO_2} + k_{O,NO_2} \cdot [O]}{k_{O_3,NO} \cdot [O_3]} \quad \text{Solve for NO to NO}_2 \text{ ratio}$$

Substitute into Ozone Continuity Equation

O₃ continuity equation for NO_x reactions

$$\left. \frac{dO_x}{dt} \right|_{NO_x} = -k_{O_3,NO} \cdot [O_3] \cdot [NO] - k_{O,NO_2} \cdot [O] \cdot [NO_2] + J(NO_2) \cdot [NO_2]$$

NO continuity equation

$$0 = -k_{NO,O_3} \cdot [O_3] \cdot [NO] + k_{O,NO_2} \cdot [O] \cdot [NO_2] + J_{NO_2} \cdot [NO_2]$$

Subtract to obtain:

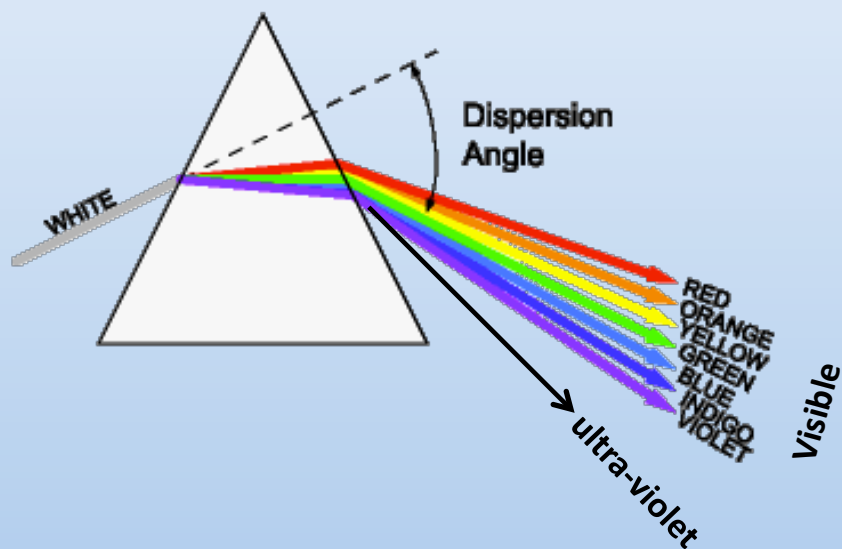
$$\left. \frac{dO_x}{dt} \right|_{NO_x} = -2 \cdot k_{O,NO_2} \cdot [O] \cdot [NO_2]$$

When added to the pure oxygen continuity equation:

$$\frac{dO_x}{dt} = 2 \cdot J(O_2) \cdot [O_2] - 2 \cdot k_{O,O_3} \cdot [O] \cdot [O_3] - 2 \cdot k_{O,NO_2} \cdot [O] \cdot [NO_2]$$

The Discovery of UV

- Johann Wilhelm Ritter uses a prism in 1801 to discover UV radiation
- Radiation beyond the violet darkened silver chloride!



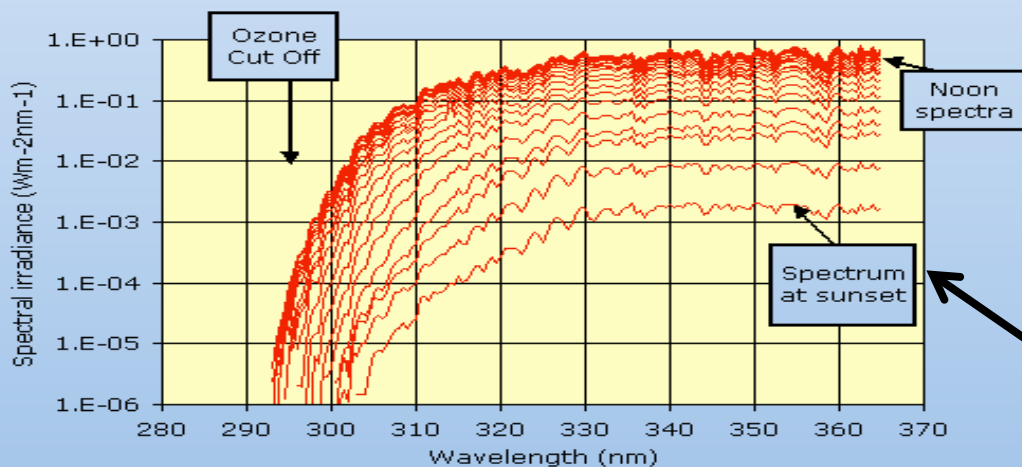
Ritter

- One year earlier Herschel had discovered infrared radiation

Cutoff of the solar UV spectrum and the discovery of atmospheric ozone

1879: Marie Alfred Cornú used newly-developed techniques for ultraviolet spectroscopy to measure the sun's spectrum.

To his surprise, the intensity of the sun's radiation dropped off rapidly at wavelengths below ~300 nm.

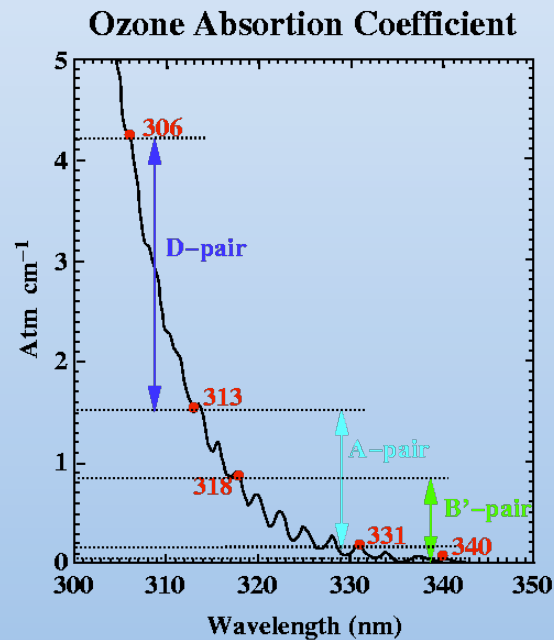
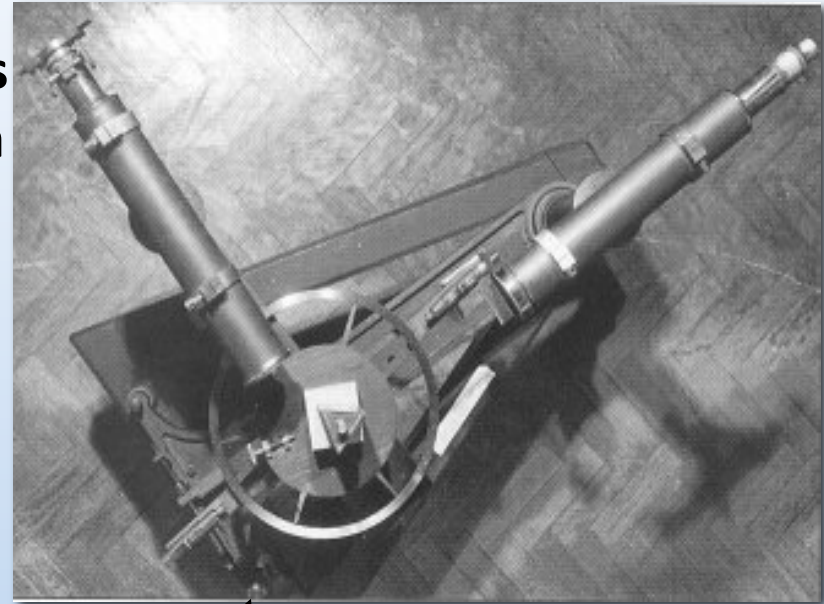


Cornú

Modern spectral measurements

Hartley and the Identification of Atmospheric Ozone

- **1883: Hartley identified ozone as the substance absorbing uv from the sun at wavelengths less than 290 nanometers**

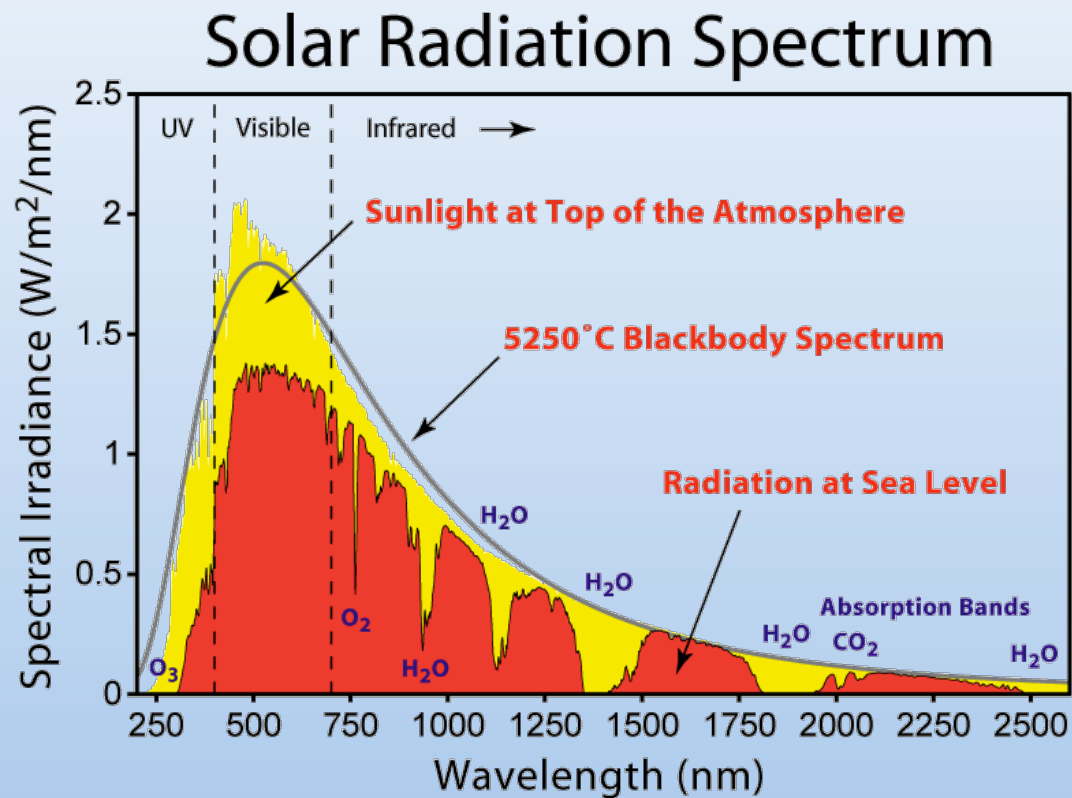


Hartley's Spectrometer
(note prism)

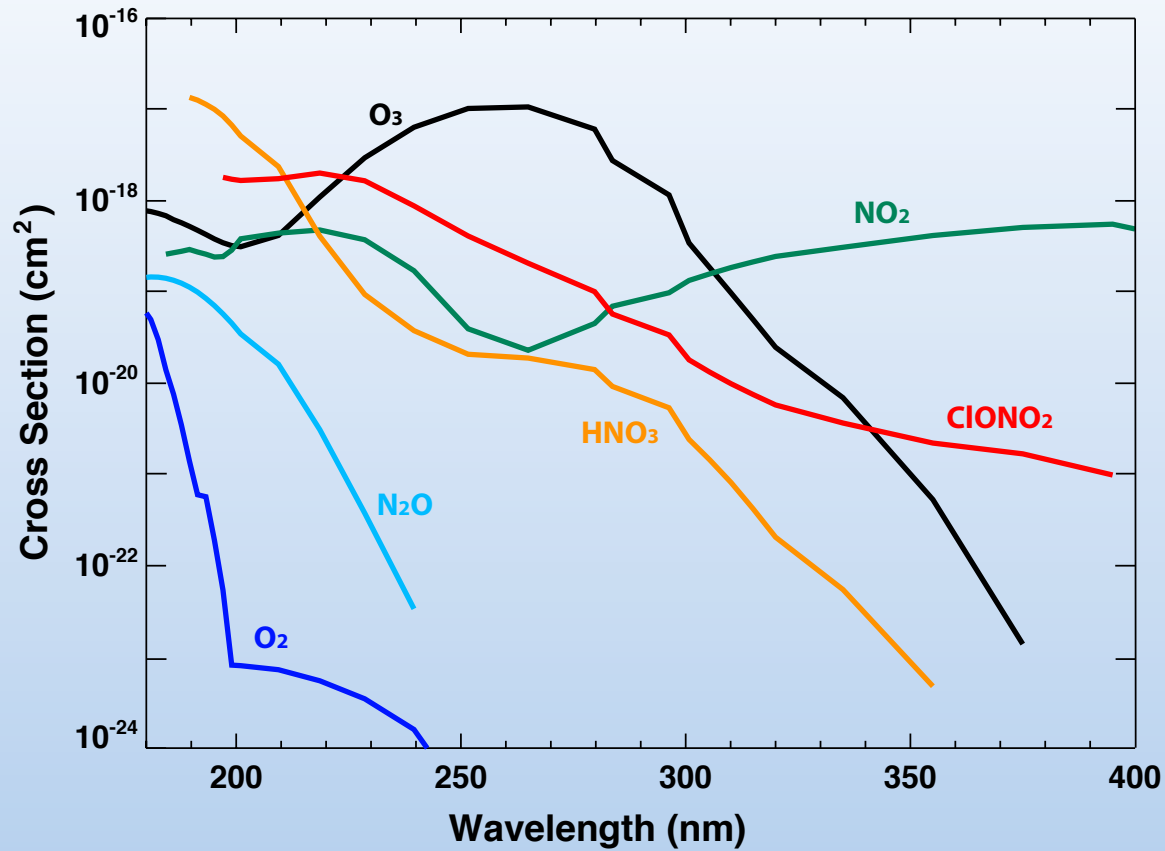


Sir Walter Noel Hartley

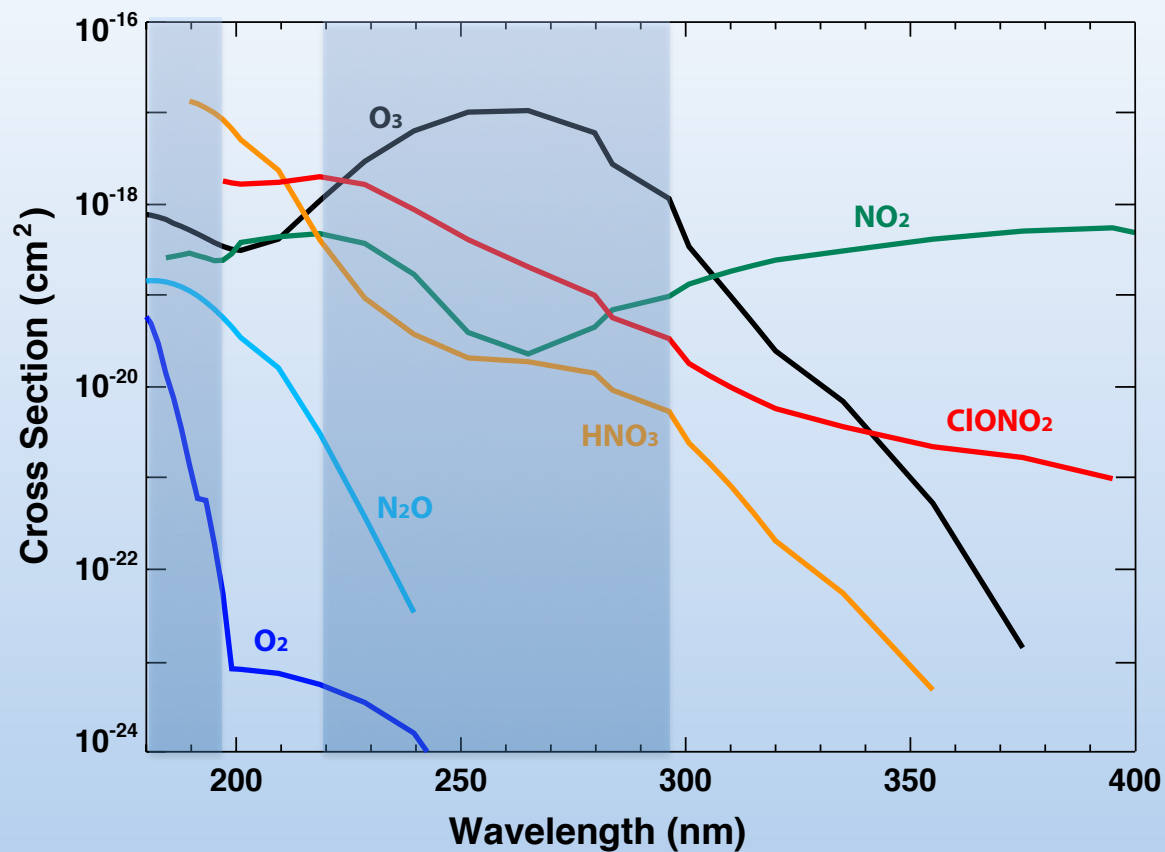
Back to Solar UV Radiation



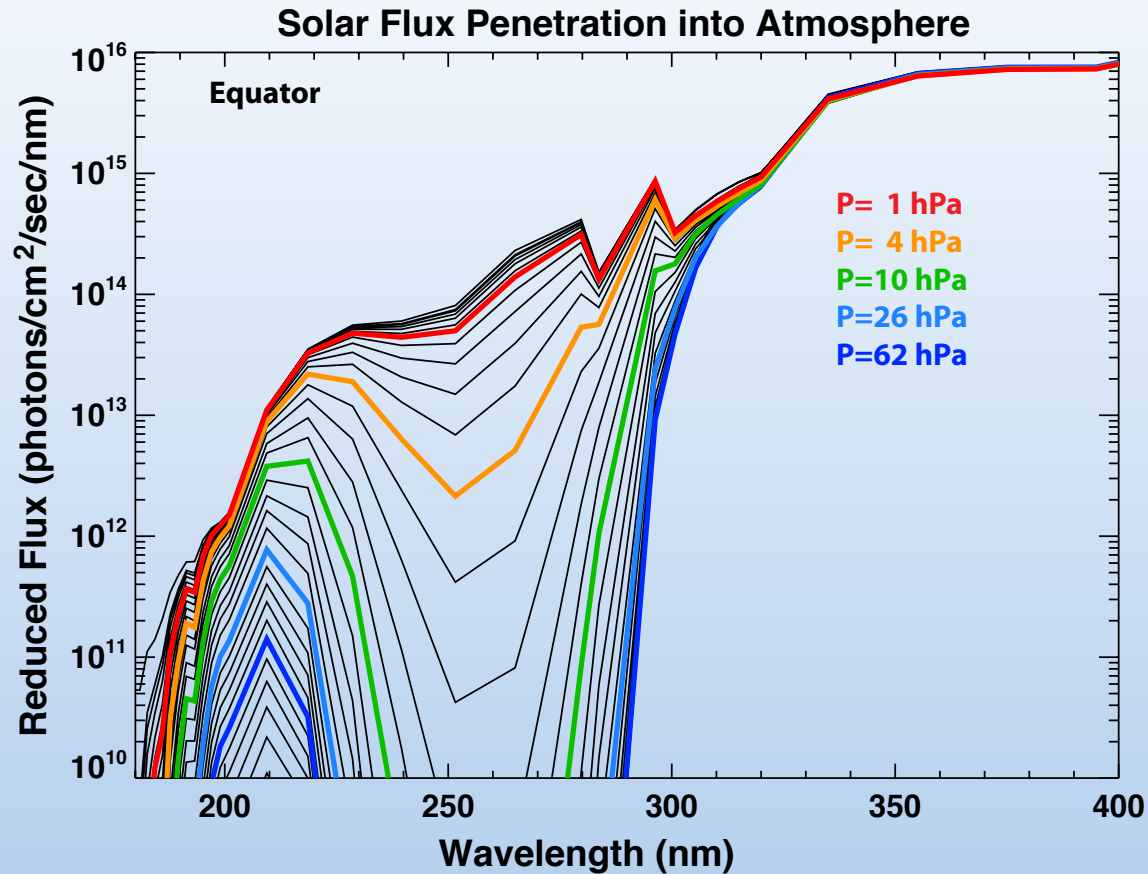
Cross Sections for Photolysis



Cross Sections for Photolysis



Reduced Solar Flux vs Altitude

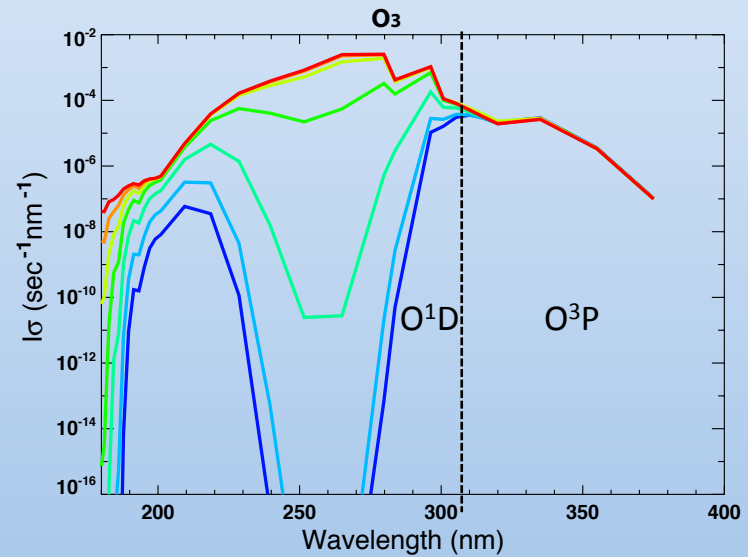
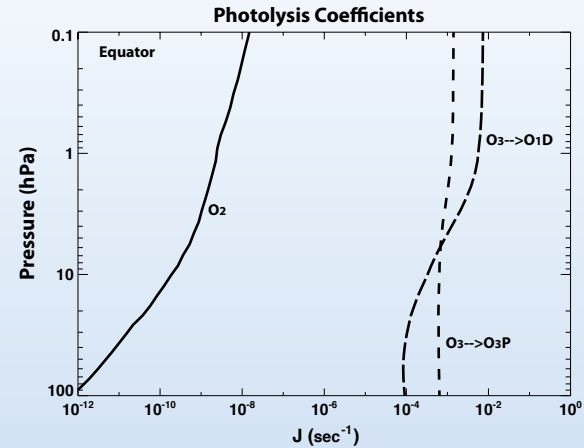
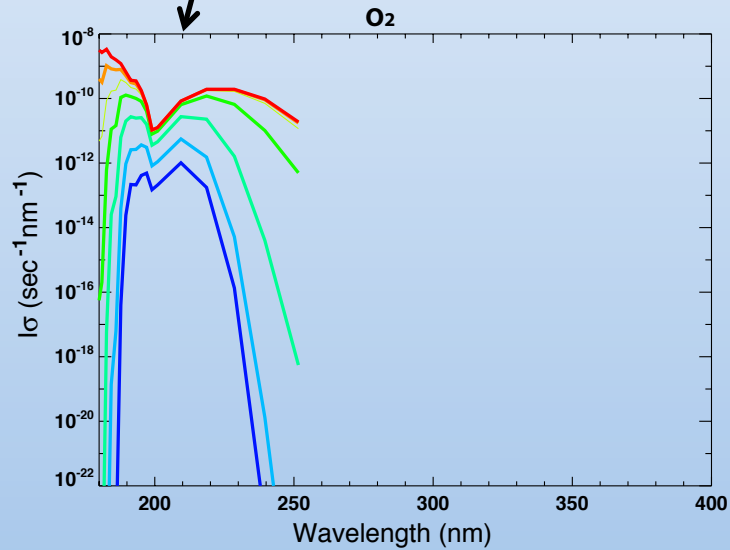


$$I(\lambda, z) = I_0(\lambda) \cdot e^{-N_{O_2}(z) \cdot \sigma_{O_2}(\lambda) - N_{O_3}(z) \cdot \sigma_{O_3}(\lambda)}$$

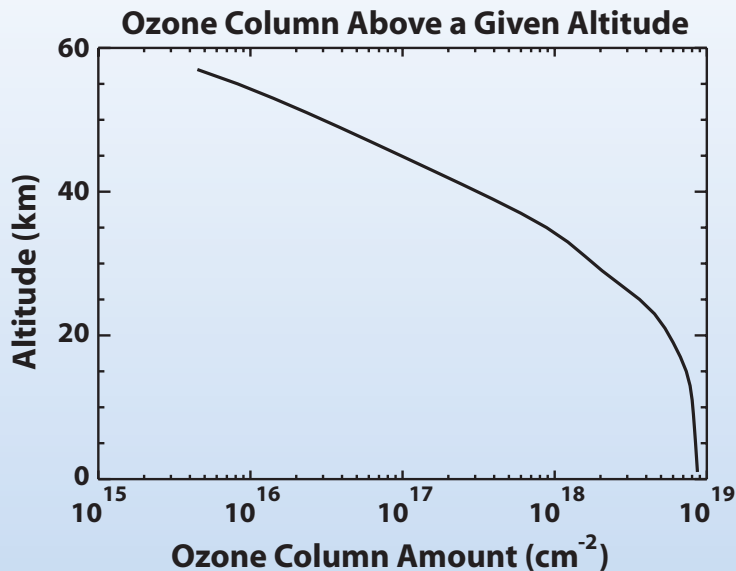
Photolysis Rates vs Wavelength

$$J_x(z) = \int \sigma_x(\lambda) \cdot I(\lambda, z) d\lambda$$

O₂ photolysis occurs mainly in the “window” near 200 nm



“Self-Healing”



Ozone Production

$$P_{O_3} = 2 \cdot J_{O_2} \cdot [O_2]$$

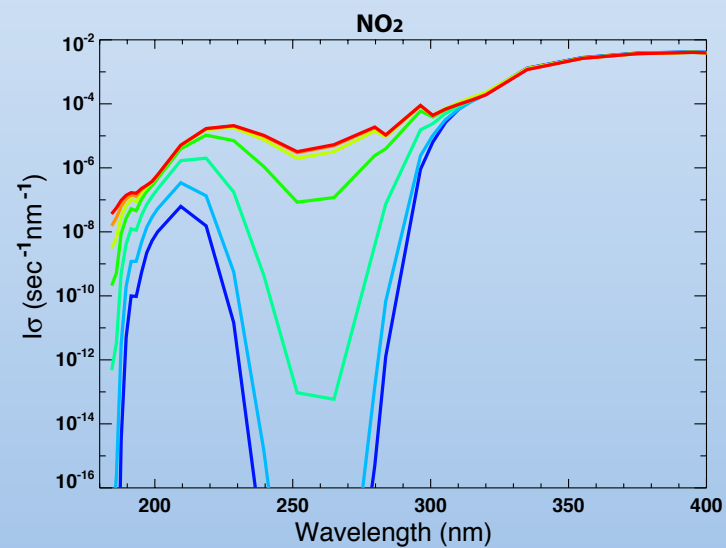
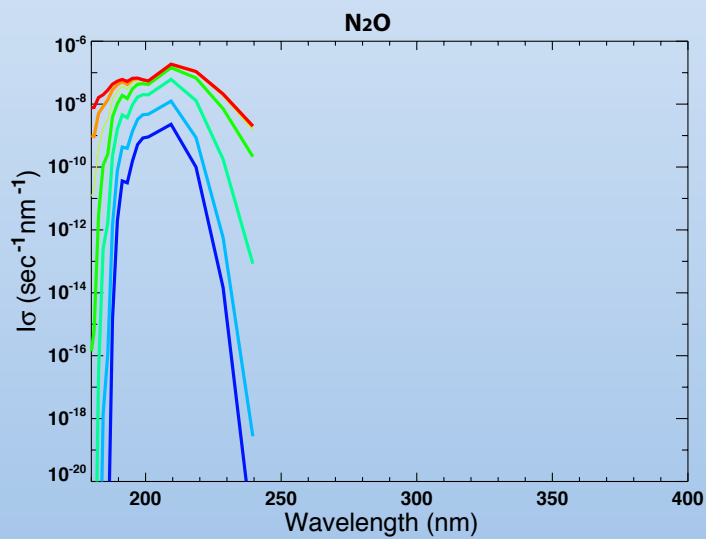
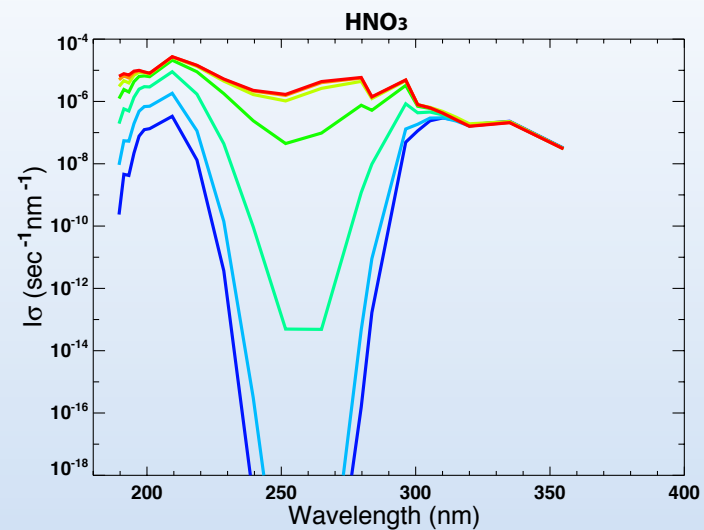
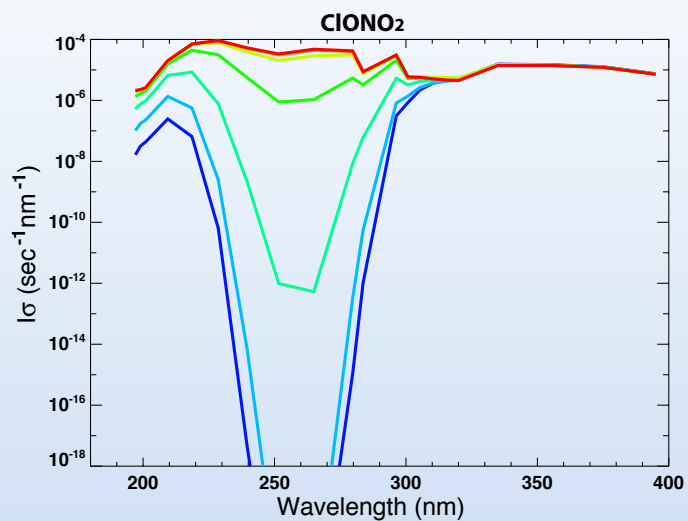
Decrease of ozone above altitude, z , results in decrease of radiative flux that leads to ozone production.



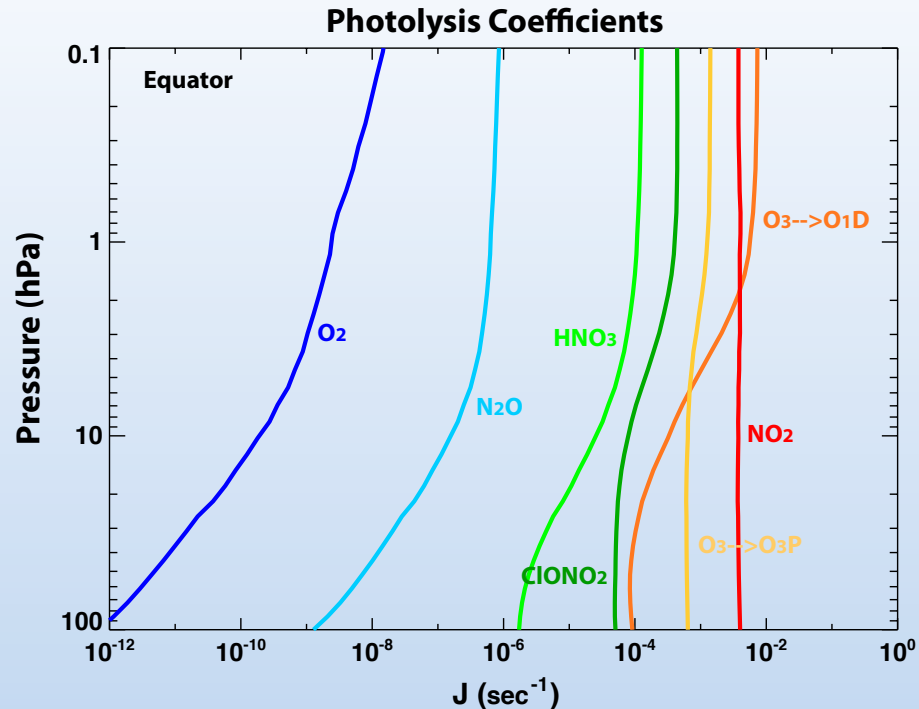
$$J_{O_2}(z) = \int \sigma_{O_2}(\lambda) \cdot I_0 e^{-[N_{O_2}(z) \cdot \sigma_{O_2}(\lambda) + N_{O_3}(z) \cdot \sigma_{O_3}(\lambda)]} d\lambda$$

Result is that a decrease in ozone above a given altitude leads to an increase in ozone at that altitude; a sort of “self-healing”.

Photolysis Rates vs Wavelength



Photolysis Coefficients vs Altitude



- Molecules that absorb in the near UV and visible have significant photolysis rates in the troposphere
- Molecules that absorb only below 300 nm have no significant photolysis in the troposphere
- Molecules that absorb only in the 200 nm window have long lifetimes unless they have chemical reactions that limit their lifetime